

A Time-Regularized, Multiple Gravity-Assist Low-Thrust, Bounded-Impulse Model for Trajectory Optimization

Donald H. Ellison ¹ Jacob A. Englander ² Bruce A. Conway ¹

¹University of Illinois at Urbana-Champaign

²Goddard Space Flight Center

February 7, 2017



Outline

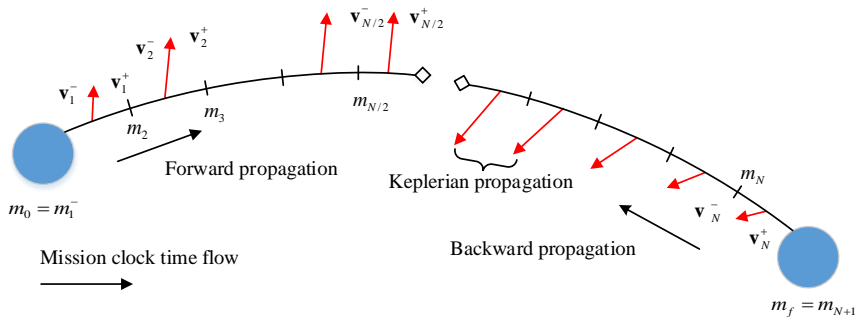
- Introduction
 - The multiple gravity-assist, low-thrust (MGALT) model
 - Time regularization
 - multiple gravity-assist, low-thrust Sundman (MGALTS) model
- Universal Anomaly Algorithm
 - The Sundman transformation
 - Extending MGALT to MGALTS
 - The Multiple Gravity Assist Low-Thrust (MGALT) model
- True Anomaly Algorithm
 - The generalized Sundman transformation
 - Extending MGALT to MGALTS
 - The Multiple Gravity Assist Low-Thrust (MGALT) model
- Analytic Derivatives
- Application: Rendezvous with Comet 45P/Honda-Mrkos-Pajdušáková
- Summary

Bounded-Impulse Trajectory Models

- Integral aspect of preliminary mission design
- Provide an accurate mass budget, if not trajectory
- EMTG, MALTO, PaGMO, GALLOP
 - Sims-Flanagan transcription, (Sims and Flanagan, 1999)
- Medium-fidelity solutions used as initial guesses for flight-qualified solutions
 - Mystic, GMAT

$$\begin{bmatrix} \mathbf{r}_{k+1} \\ \mathbf{v}_{k+1} \end{bmatrix} = \begin{bmatrix} F & G \\ \dot{F} & \dot{G} \end{bmatrix} \begin{bmatrix} \mathbf{r}_k \\ \mathbf{v}_k \end{bmatrix} \quad (1)$$

Multiple Gravity-Assist Low-Thrust Model



- Sims-Flanagan and patched conic approximation, phase divided into N control segments
- Control nodes evenly spaced in time $\Delta t = \frac{\Delta t_p}{N}$

Modeling High-Eccentricity with MGALT

- Equal temporal spacing of control nodes is non-optimal for increasing eccentricity.
- Apoapse raising is most efficiently achieved via maneuvering at periapse.
- Using additional segments results in an increase in problem dimension.
- Redistribution of the same number of nodes would be more desirable.
- This is typically achieved with a form of time regularization.

The Sundman Transformation

- Differential transformation introduced by Karl Sundman to regularize the three-body problem.
- A new variable χ replaces time as the independent propagation variable:

$$dt = \frac{r}{\sqrt{\mu}} d\chi \quad (2)$$

- This results in an even geometric distribution of control nodes around the orbit.

Multiple Gravity-Assist Low-Thrust Sundman

- The MGALT model encodes the phase flight time Δt_p as a decision variable
- MGALTS adds a total angular propagation variable $\Delta\chi_p$
- Each segment is propagated for a portion of the total independent variable:
$$\Delta\chi = \frac{\Delta\chi_p}{N}$$
- The propagation time associated with each $\Delta\chi$ is computed using a universal Kepler solver in an inverse sense.
 - χ is transcendental in Kepler's equation, but is now selected by NLP solver
 \implies iteration is no longer required in Kepler solver
- An additional nonlinear match point constraint is required to ensure match point epoch is continuous:

$$\mathbf{c}^\dagger = \begin{bmatrix} \mathbf{r}_B - \mathbf{r}_F \\ \mathbf{v}_B - \mathbf{v}_F \\ m_B - m_F \\ t_B - t_F \end{bmatrix} = \mathbf{0} \quad (3)$$

- Analytic derivatives of the match point constraint are still possible to compute

Calculating Maximum Impulse Magnitudes

- For MGALT, the maximum permissible Δv in a segment is computed as follows:

$$\Delta v_{\max} = \frac{N_{\text{active}} D T_{\max} \Delta t}{m_k} \quad (4)$$

- For MGALTS, this can only be computed after the segment propagation time Δt has been determined
- Impulses shifted from the center of the segments to the boundary
- The maneuver at the match point can *potentially* be twice the size of the others
 - This potential reduces as the number of segments increases.

True Anomaly as the Independent Variable

- In order to achieve a different control node distribution, consider the generalized Sundman transformation

$$dt = c r^\gamma d\chi \quad c \in \mathbb{R} \quad \text{and} \quad \gamma \in \{\mathbb{R} : \gamma \geq 0\} \quad (5)$$

- Universal Kepler propagator based on $\{U_n(\chi; \alpha)\}$, derivations based on $\gamma = 1$ corresponding to the eccentric anomaly.
- Require a relationship between Δf and $\Delta\chi$

$$\Delta\chi = \begin{cases} \frac{2r_k \sin(\frac{1}{2}\Delta f)}{\sqrt{p} \cos(\frac{1}{2}\Delta f) - \sigma_k \sin(\frac{1}{2}\Delta f)} & ; \alpha = 0 \\ \frac{2}{\sqrt{\alpha}} \tan^{-1} \left[\frac{\sqrt{\alpha} r_k \sin(\frac{1}{2}\Delta f)}{\sqrt{p} \cos(\frac{1}{2}\Delta f) - \sigma_k \sin(\frac{1}{2}\Delta f)} \right] & ; \alpha > 0 \\ \frac{2}{\sqrt{-\alpha}} \tanh^{-1} \left[\frac{\sqrt{-\alpha} r_k \sin(\frac{1}{2}\Delta f)}{\sqrt{p} \cos(\frac{1}{2}\Delta f) - \sigma_k \sin(\frac{1}{2}\Delta f)} \right] & ; \alpha < 0 \end{cases} \quad (6)$$

Rendezvous with Comet 45P/Honda-Mrkos-Pajdušáková

Parameter	Value
Maximum allowed initial mass	4000 kg
Launch window	5 years
Maximum flight time	unbounded
Earliest allowed launch date	August 27 th , 2016
Latest allowed arrival date	January 1 st , 2030
Launch vehicle	Atlas V (555) NLS-2
Maximum launch C3	21.7156 km ² /s ²
Launch declination bounds	[−28.5°, 28.5°]
Thruster	2 × NEXT TT11 high thrust
Throttle logic	minimum number of thrusters
Thruster duty cycle	0.95
Solar array BOL power (launch at 1 A.U.)	40.0 kW
Power margin (δ_{power})	15%
Post-launch checkout coast	60 days
Number of segments per phase	200
Ephemeris	SPICE
Objective function	maximize final mass

- Solar electric propulsion model adheres to Discovery class mission proposal requirements

Rendezvous with Comet 45P: MGALT

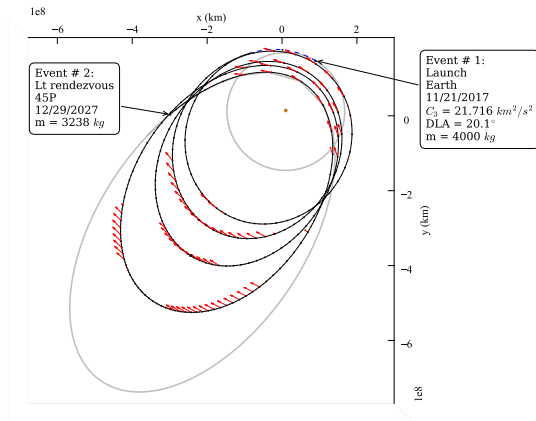


Figure: Rendezvous with 45P using MGALT with 200 segments, $\Delta t_{\text{mission}} = 3690$ days

Rendezvous with Comet 45P: MGALTS

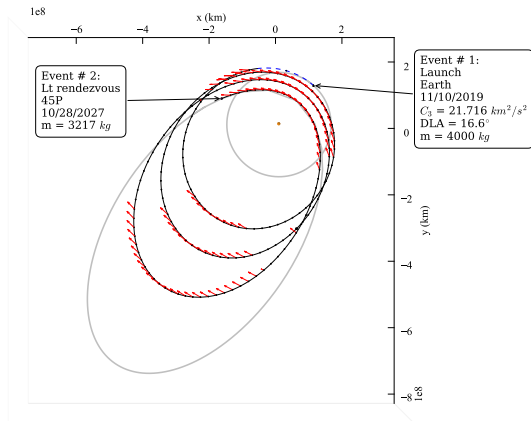
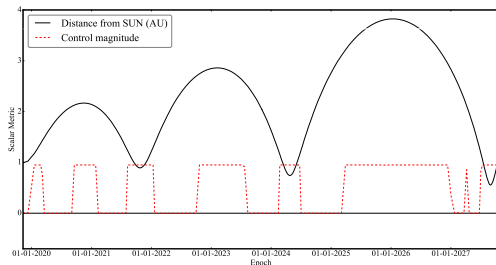
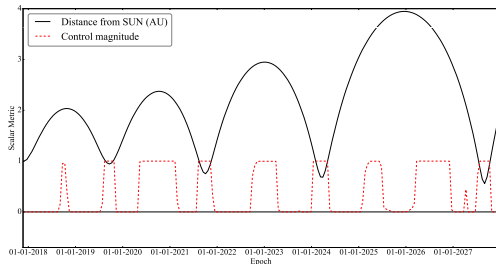


Figure: Rendezvous with 45P using MGALTS with 200 segments, $\Delta t_{\text{mission}} = 2909$ days

Rendezvous with Comet 45P: Comparison



- MGALT requires an additional revolution of the sun
- Additional nodes at periapse affords MGALTS more control authority
- MGALTS TOF is substantially less

Rendezvous with Comet 45P: MGALT More Nodes!

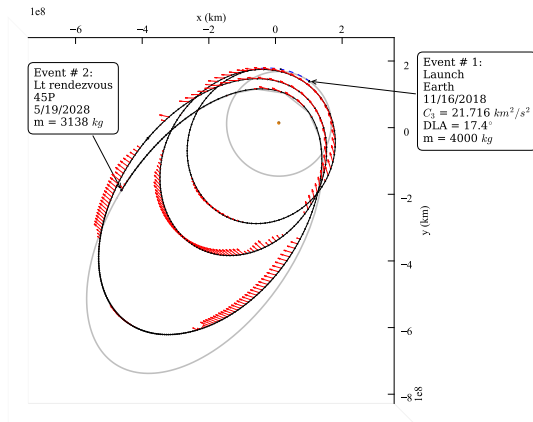


Figure: Rendezvous using MGALT with 400 segments, $\Delta t_{\text{mission}} = 3472$ days

- MGALTS can achieve the same transfer using far fewer control nodes

Rendezvous with Comet 45P: MGALTS True Anomaly

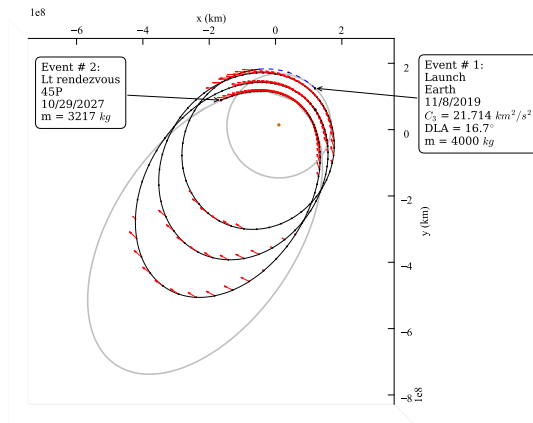
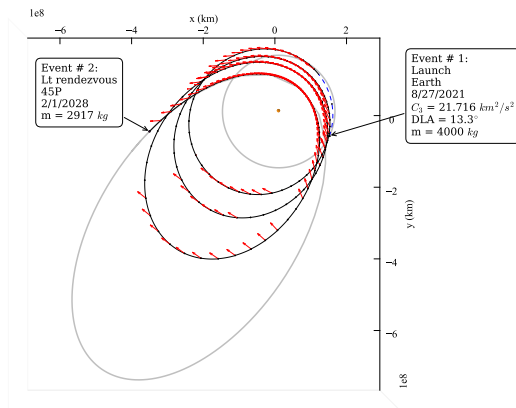


Figure: Rendezvous using the MGALTS true anomaly-based transcription with 200 segments, $\Delta t_{\text{mission}} = 2913$ days

Rendezvous with Comet 45P: MGALTS True Anomaly 2



- Using true anomaly as the independent variable, nodes are now concentrated at periapse
- Short flight time pump-up maneuvers are accessible
- Useful inside an outer loop that is capable of exploring wide flight time ranges

Figure: MGALTS true anomaly, 200 segments,
 $\Delta t_{\text{mission}} = 2350 \text{ days}$

Summary

- A medium-fidelity low-thrust model has been developed that solves the time-regularized Kepler problem exactly.
- Straightforward extension of the MGALT model in use by many preliminary design tools
- Kepler propagation is made iteration free at the expense of one additional decision variable and constraint per phase.
- Especially useful for trajectories requiring large eccentricity increases
 - e.g. Pump-up/down maneuvers
- MGALTS can perform certain transfers using fewer control nodes than MGALT.
- Certain solution families can only reasonably be accessed using MGALTS.

Thank You

- Ricardo Restrepo and Ryan Russell for inspiring the iteration-free Kepler propagator.
- Kyle Hughes and Jeremy Knittel (GSFC 595) for reviewing.

